"...an excellent textbook, with a clear and pedagogical presentation. The additions in this second edition, such as galaxy formation and baryon acoustic oscillations, are valuable and bring the book even more up-to-date."

Professor Zoltán Haiman, Columbia University, New York

"Barbara Ryden writes in a very clear and engaging style. This transparency has inspired many undergraduate science majors in my cosmology class to pursue additional coursework and research in astrophysics."

Professor Crystal Martin, University of California, Santa Barbara

"[This] book paints an elegant mathematical picture of the evolution of the universe from the Big Bang to the formation of stars. Ryden does a masterful job of paring cosmology down to its most fundamental elements and presenting complex topics with exceptional clarity."

Christy Tremonti, University of Wisconsin, Madison

The second edition of *Introduction to Cosmology* is an exciting update of this award-winning textbook. It is aimed primarily at advanced undergraduate students in physics and astronomy, but is also useful as a supplementary text at higher levels. It explains modern cosmological concepts, such as dark energy, in the context of the Big Bang theory. Its clear, lucid writing style, with a wealth of useful everyday analogies, makes it exceptionally engaging. Emphasis is placed on the links between theoretical concepts of cosmology and the observable properties of the universe, building deeper physical insights in the reader. The second edition includes:

- up-to-date observational results;
- fuller descriptions of special and general relativity;
- expanded discussions of dark energy;
- a new chapter on baryonic matter that makes up stars and galaxies.

It is an ideal textbook for the era of precision cosmology in the accelerating universe.



For instructors use:

▶ PowerPoint slides and JPEGS of figures and tables from the book

¥ Solutions manual available as PDF

Cover image: Large scale projection through the Illustris volume at z=0, centered on the most massive cluster, 15 Mpc/h deep. It shows dark matter density (left) transitioning to gas density (right) [Credit: Illustris Collaboration]

C UN ww

CAMBRIDGE UNIVERSITY PRESS www.cambridge.org



	Pref	face Symbols S	page xi
1	Intro	oduction Demerken	1
2	Fund	lamental Observations	6
	2.1	The Night Sky is Dark	6
	2.2	The Universe is Isotropic and Homogeneous	9
	2.3	Redshift is Proportional to Distance	12
	2.4	Different Types of Particles	18
	2.5	Cosmic Microwave Background	23
	Exer	rcises	25
3	New	ton versus Einstein	27
	3.1	The Way of Newton	28
	3.2	The Special Way of Einstein	29
	3.3	The General Way of Einstein	34
	3.4	Describing Curvature	37
	3.5	The Robertson–Walker Metric	41
	3.6	Proper Distance	43
	Exer	cises	47
4	Cosn	nic Dynamics	49
	4.1	Einstein's Field Equation	50
	4.2	The Friedmann Equation	52
	4.3	The Fluid and Acceleration Equations	58

	4.4	Equations of State	60
	4.5	Learning to Love Lambda	63
	Exer	cises	67
5	Mode	el Universes	69
	5.1	Evolution of Energy Density	69
	5.2	Empty Universes	74
	5.3	Single-component Universes	77
		5.3.1 Matter only	80
		5.3.2 Radiation only	81
		5.3.3 Lambda only	- 83
	5.4	5.4 Multiple-component Universes 8	
		5.4.1 Matter + Curvature	86
		5.4.2 Matter + Lambda	90
		5.4.3 Matter + Curvature + Lambda	92
		5.4.4 Radiation + Matter	95
	5.5	Benchmark Model	96
	Exer	cises	100
6 Measuring Cosmological Parameters			102
	6.1	"A Search for Two Numbers"	102
	6.2	Luminosity Distance	106
	6.3	Angular-diameter Distance	110
	6.4	Standard Candles and H_0	114
	6.5	Standard Candles and Acceleration	116
	Exer	cises	121
7	Dark	Matter	123
•	7.1	Visible Matter	123
	7.2	Dark Matter in Galaxies	128
	7.3	Dark Matter in Clusters	130
	7.4	Gravitational Lensing	135
	7.5	What's the Matter?	139
		cises	nonumpel manufactule dT 140
			The later of the bound of the bound

1/

8	The C	Cosmic Microwave Background	142
	8.1	Observing the CMB	143
	8.2	Recombination and Decoupling	147
	8.3	The Physics of Recombination	150
	8.4	Temperature Fluctuations	157
	8.5	What Causes the Fluctuations?	159
	Exerc	cises	164
9	Nucle	eosynthesis and the Early Universe	166
	9.1	Nuclear Physics and Cosmology	167
	9.2	Neutrons and Protons	169
	9.3	Deuterium Synthesis	174
	9.4	Beyond Deuterium	177
	9.5	Baryon–Antibaryon Asymmetry	181
	Exer	cises	183
10 Inflation and the Very Early Universe			
	10.1	The Flatness Problem	186
	10.2	The Horizon Problem	187
	10.3	The Monopole Problem	189
	10.4	The Inflation Solution	192
	10.5	The Physics of Inflation	197
	Exer	cises	202
П	Struc	cture Formation: Gravitational Instability	204
	11.1	The Matthew Effect	206
	11.2	The Jeans Length	209
	11.3	Instability in an Expanding Universe	213
	11.4	The Power Spectrum	217
	11.5	Hot versus Cold	221
	11.6	Baryon Acoustic Oscillations	226
	Exer	cises	230

12 Structure Formation: Baryons and Photons 232				
12.1 Baryonic Matter Today	233			
12.2 Reionization of Hydrogen	235			
12.3 The First Stars and Quasars	238			
12.4 Making Galaxies	242			
12.5 Making Stars	Yemothanisali ada eegaa 248			
Exercises	254			
Epilogue	256			
Table of Useful Constants Index	258 259			